Common Parameters

The Attribute Manager will show the following common attributes under the Main tab group:
**Color**

The color of the light.

**Intensity**

Intensity controls the brightness of light emitted by the light source by multiplying the color.

**Exposure**

Exposure is an f-stop value that multiplies the intensity by 2 to the power of the f-stop. Increasing the exposure by 1 results in double the amount of light.

In Arnold, the total intensity of the light is computed with the following formula:

\[
\text{color} \times \text{intensity} \times 2^{\text{exposure}}
\]

You can get the same output by modifying either the intensity or the exposure. For example, intensity=1, exposure=4 is the same as intensity=16, exposure=0. **Note:** \(2^0 = 1\), not 0.

\[
1 \times 1 \times 2^4 = 16
\]

\[
1 \times 16 \times 2^0 = 16
\]

The reasoning behind this apparent redundancy is that, for some people, f-stops are a much more intuitive way of describing light brightness than raw intensity values, especially when you're directly matching values to a plate. You may be asked by the director of photography (who is used to working with camera f-stop values) to increase or decrease a certain light by 'one-stop'. Other than that, this light parameter has nothing to do with a real camera's f-stop control. Also, working with exposure means you won't have to type in huge values like 10,000 in the intensity input if your lights have quadratic falloff (which they should).

If you are not used to working with exposure in the lights, you can simply leave the exposure parameter at its default value of 0 (since \(2^0 = 1\), the formula then simplifies to: color \(\times\) intensity \(\times\) 1).

**Samples**

Controls the quality of the noise in the soft shadows and direct specular highlight. The higher the number of samples, the lower the noise, and the longer it takes to render. The exact number of shadow rays sent to the light is the square of this value multiplied by the AA samples.
Noise from lights can sometimes be difficult to diagnose, particularly if the light source is broad in comparison to the scene and the shadows have an extremely wide penumbra. In these cases, it can sometimes be mistaken for indirect diffuse noise. It highlights the necessity for testing noise ray type. This diagram shows how light is traced in Arnold.

The example below shows specular highlights from area light sources. There are four spherical mesh light sources of varying sizes and color temperatures. Underneath are four cubes with Standard shaders assigned to them with varying degrees of specular roughness. Note that more noise is apparent with smaller light sources. Increasing the number of light samples resolves the noise.
If the issue is noise in a specular highlight, you will need to confirm that the source is the direct light and not a secondary ray type (such as specular). This is easy to achieve by turning off global illumination; set the **Diffuse Depth** and **Specular Depth** to zero (this essentially turns off all global illumination). If the noise is still there, we know it is the specular component of the illumination model. If the issue is shadow noise, then we can simply toggle ignore shadows in the Arnold render settings, and the noise will completely resolve.

The key, again, is to modify the sampling and observe the changes. Increasing the number of light samples should have an immediate and quantitative effect on the smoothness of the specular highlight and the shadow. If there is no change, light samples are not responsible for the noise.

Multiple importance sampling (MIS) is enabled by default in Arnold. The images below show the difference when rendering with and without multiple importance sampling.

More information about light samples and removing noise can be found [here](#).
**Normalize**

If enabled, you will be able to tweak the shadow softness by changing the size (i.e., radius) of the light, without affecting the amount of emitted light. This is very handy for artistic control. Otherwise, if not enabled, the amount of emitted light is proportional to the light's surface area.

Enabled (default). Radius 2.5.  
Disabled. Radius 2.5.

---

**Cast Shadows**

Enables the computation of shadows cast from the light.

Enabled (default)  
Disabled

---

**Shadow Density**

Sets the shadow density, or strength. This controls how the shadow blends with the material on which the shadow is cast: a value of 1.0 produces an opaque, black shadow, and a value of 0.0 gives no shadow. Normally this would be 1.0.
**Shadow Color**

Sets the intensity of each color channel for shadows. Normally this would be black.

- **Black (default)**
- **Red**

**Cast Volumetric Shadows**

Determines if volumetric shadows will be computed. This option is not in Distant or Skydome Lights.

- **Enabled (default)**
- **Disabled**

**Volume Samples**
The volume samples parameter handles the number of samples used to integrate the in-scattering from direct light. Like the "samples" parameter for surfaces, it is also a squared number.

**Volume Samples** do not apply to the atmosphere shader (e.g., fog or atmosphere_volume) in the Environment section. To improve the quality of the atmosphere shader, you must increase the atmosphere_volume samples.

**Light Contribution**

Per-light scaling for Camera, Transmission, Diffuse, Specular, SSS, Indirect and Volume. Weights scaling the light contribution to each of those components independently. Should be left at 1 to produce physically accurate results.

Only Area lights and Point lights (non-0 radius) are visible to the camera. Camera and Transmission values default to 0 with Area lights.

**Indirect**

The relative energy loss (or gain) at each bounce. This should be left at its default value of 1.0 for physically meaningful results. Values bigger than 1 make it impossible for GI algorithms to converge to a stable solution.
**Max Bounces**

The maximum number of times the energy from this light will be allowed to bounce in the scene. A Max Bounces value of 0 means that the light will only be part of the direct lighting computations, effectively disabling GI for this light. Note that this value works along with the global ray depth controls, so the default value of 999 bounces per light is just a theoretical maximum; in practice, the global ray depth limits are set much lower.

![999 (default)](image1) ![0](image2)

**AOV Light Group**

Per-light AOVs are available via a string parameter. Each light has an AOV parameter which writes out the light contribution to a separate AOV with a corresponding name. See also Light Path Expression AOVs.

![Volume Beauty AOV](image3) ![Red light volume AOV](image4)

Information about creating per-light AOVs can be found here.
Details Tab

Use Color Temperature

The temperature of an ideal black-body radiator, in kelvin units, that is used to determine the color for a light source. The default color is set to 6500 K, which is considered as the white point by the Commission Internationale de l'Eclairage (CIE). The color ranges from red, through to white and then to blue. Values above 6500 K will give a cool color, whilst values below will show a warm color.

The image above shows the effect of color temperatures on a scene in kelvin units.
A range of color temperature values in kelvin (in this case a cylinder light has been used).

Note that **Use Color Temperature** will override the default color for the light. That includes any textures that are assigned to the color attribute.

![Texture map connected to color attribute of light](image1)
![Color temperature enabled overrides color texture map](image2)

**Blend**

Multiples the light color with the color temperature when enabled. Otherwise, the color temperature overwrites the light color.

**Light Filters**

Light filters can extend light sources with additional effects such as modifying a light's output based on distance, position or other factors. **Light filters** are shaders in Arnold. You can define a list of light filters by dropping them onto this field.

**User Options**

This string field can be set to override any parameters of an Arnold node or define custom parameters. See more details [here](#).